

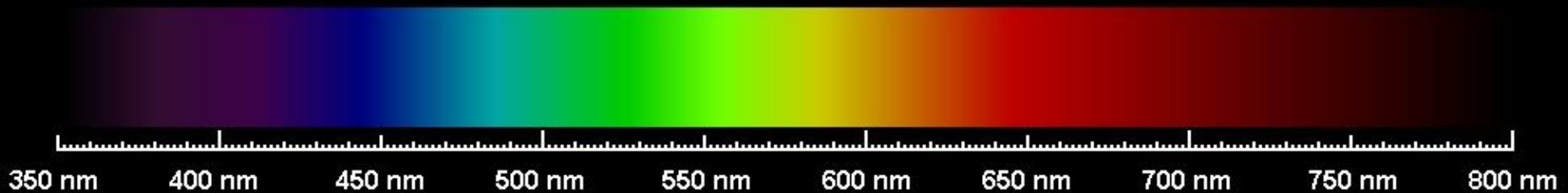
Cardiff RT 2019

A night photograph of a cityscape, likely Cardiff, with a prominent red light source in the distance. The city lights are visible in the foreground, and the red light is a bright, circular spot in the middle ground. The sky is dark, and the overall scene is illuminated by the city lights and the red light.

Communicating with Light

Barry Chambers G8AGN

The Visible Spectrum

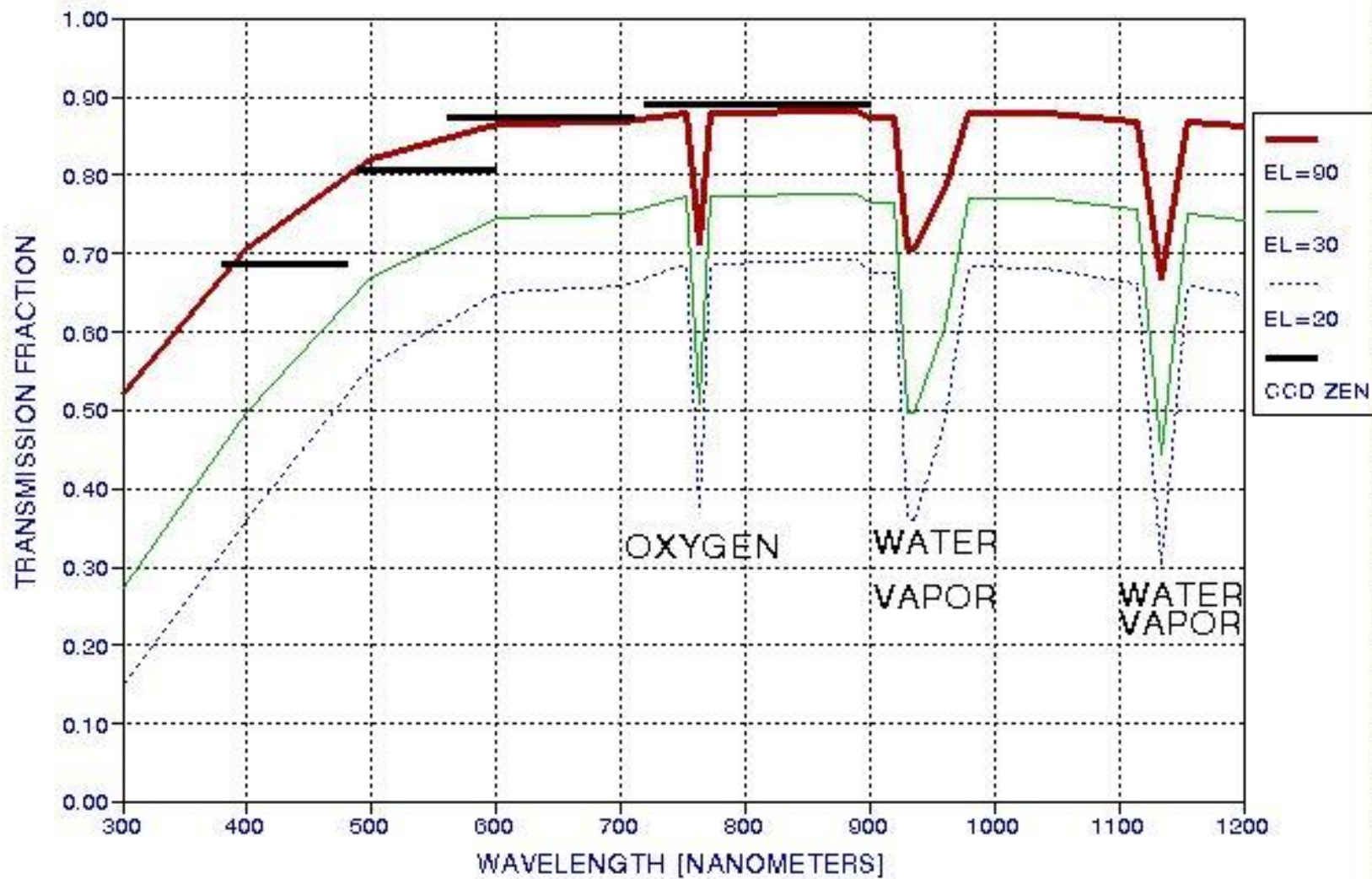


Visible Continuous Spectrum 2

(Perceived Brightness Partially to Scale)

- Red LEDs and lasers emit at wavelengths between 620–670 nm
- The eye is most sensitive to wavelengths around 550 nm
- Atmospheric transparency depends on wavelength

ATMOSPHERIC TRANSMISSION SPECTRUM TYPICAL CONDITIONS

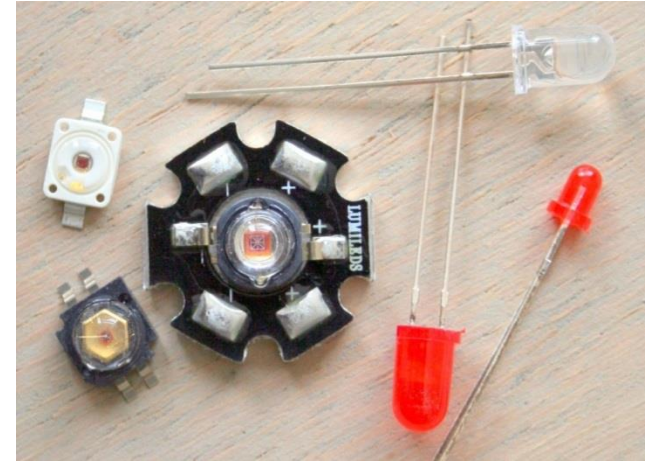
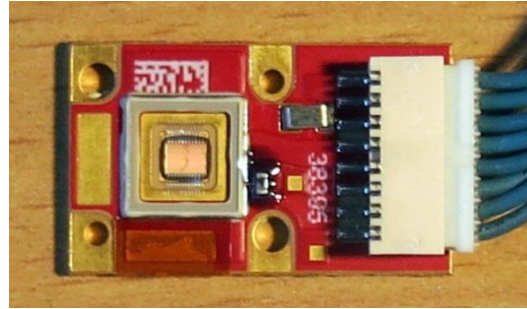


Source: B.L. GARY, 1990

Red optical sources

Luminus PhlatLight
2.5A/mm²

8A standing current
>16A peak current



LEDs

High outputs available (watts!)

Easily modulated

Wide beam (collimation)

Lasers

Very narrow beam

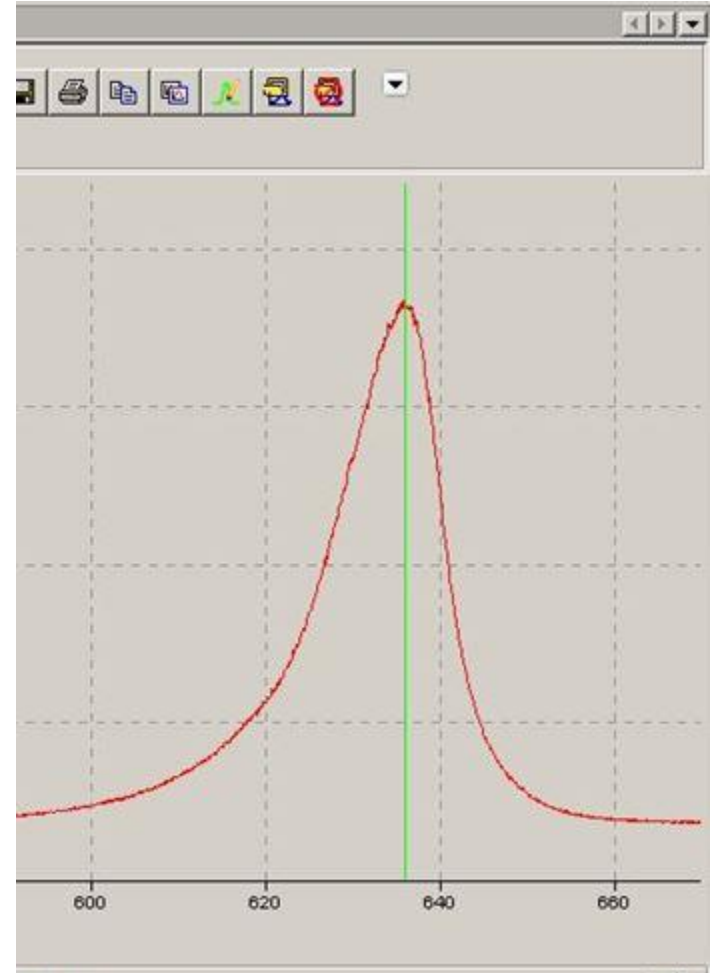
Eye safety issues



Measured spectra of red optical sources



Laser pointer



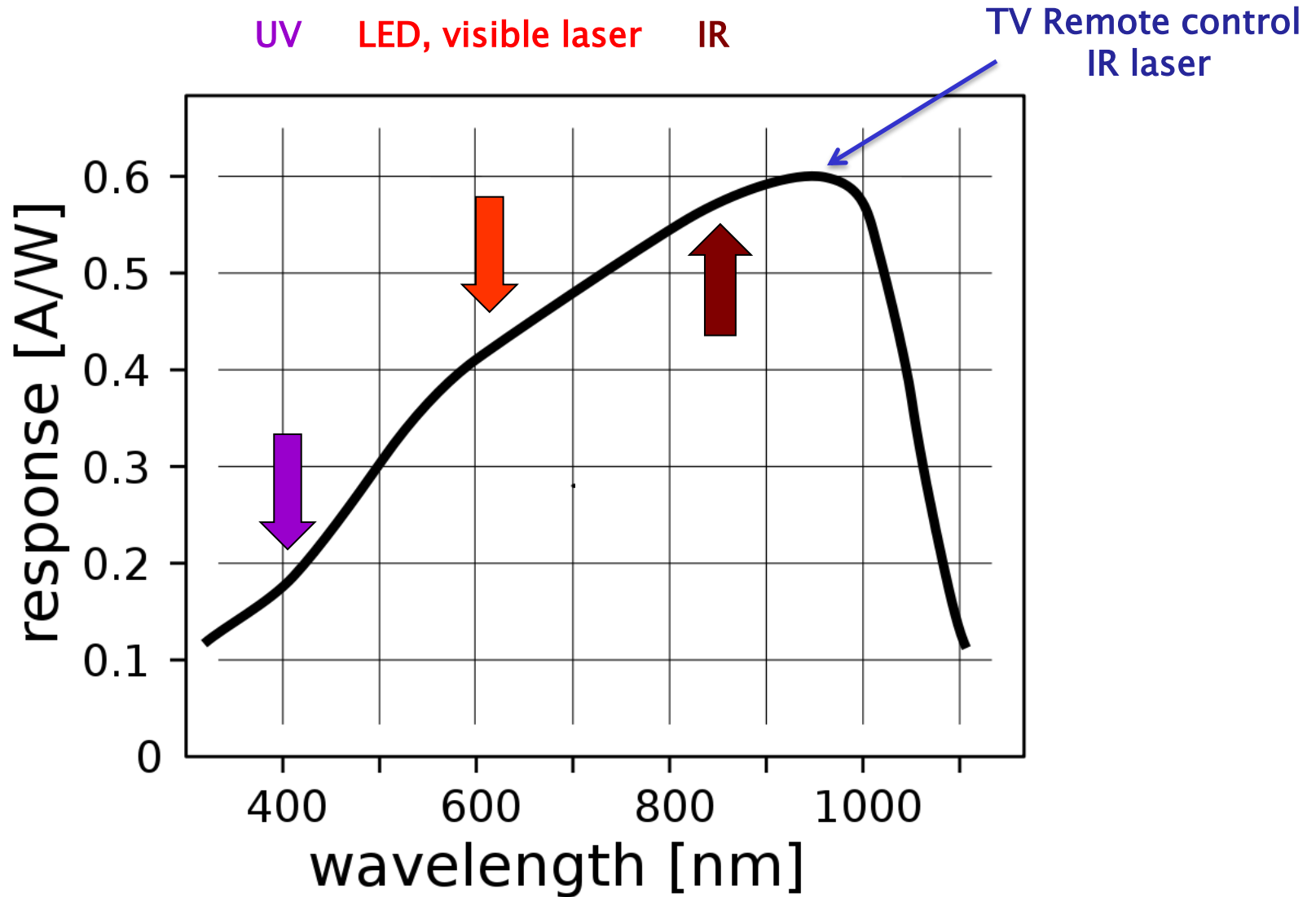
3 watt Luxeon Star

Measured with Ocean Optics HR-2000 spectrometer

Optical Detectors

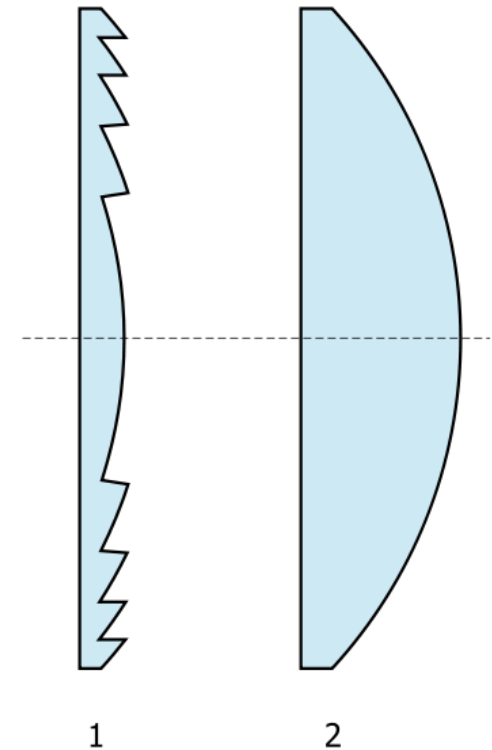
- Photo transistor
Low MHz bandwidth
Sensitive but slow & noisy
- PIN diode
Fast but lower output
Low noise – GHz bandwidth possible
- LED
High efficiency LEDs look similar to PIN diodes in sensitivity
Relative speed & noise unknown
- Avalanche photo diode
High gain, low noise, very expensive.
Need high applied reverse bias voltage

Response of Silicon photo diodes and transistors



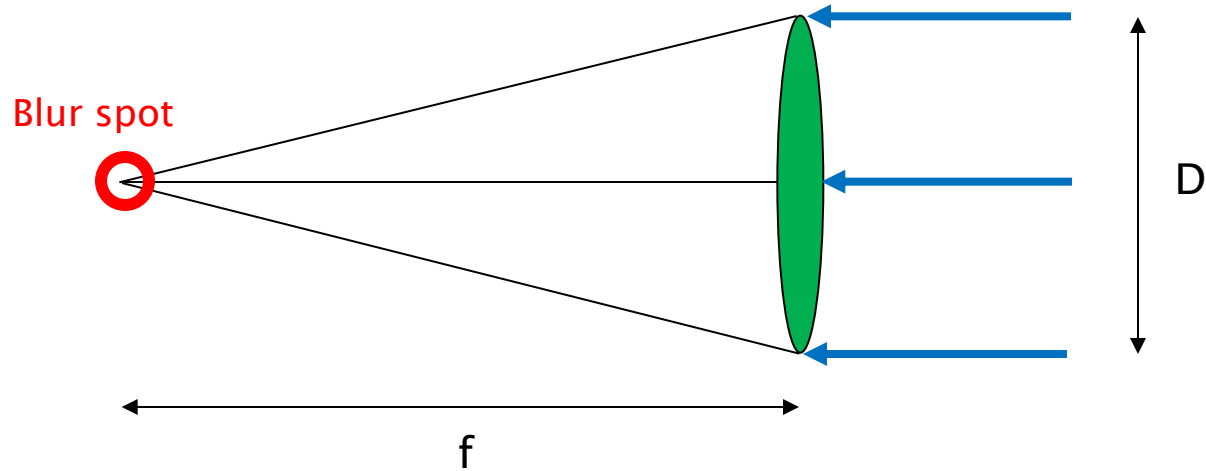
Lenses

- Act like antennas – larger aperture means higher gain with narrower beam-width
- Typical gain is approximately the ratio of lens diameter to the photo-detector area
Photodiodes are typically 1 – 3mm square
- Glass lenses are best but heavy and expensive. Plastic A4 size Fresnel lenses are OK and only 99p



1. Fresnel lens
2. Plano convex lens

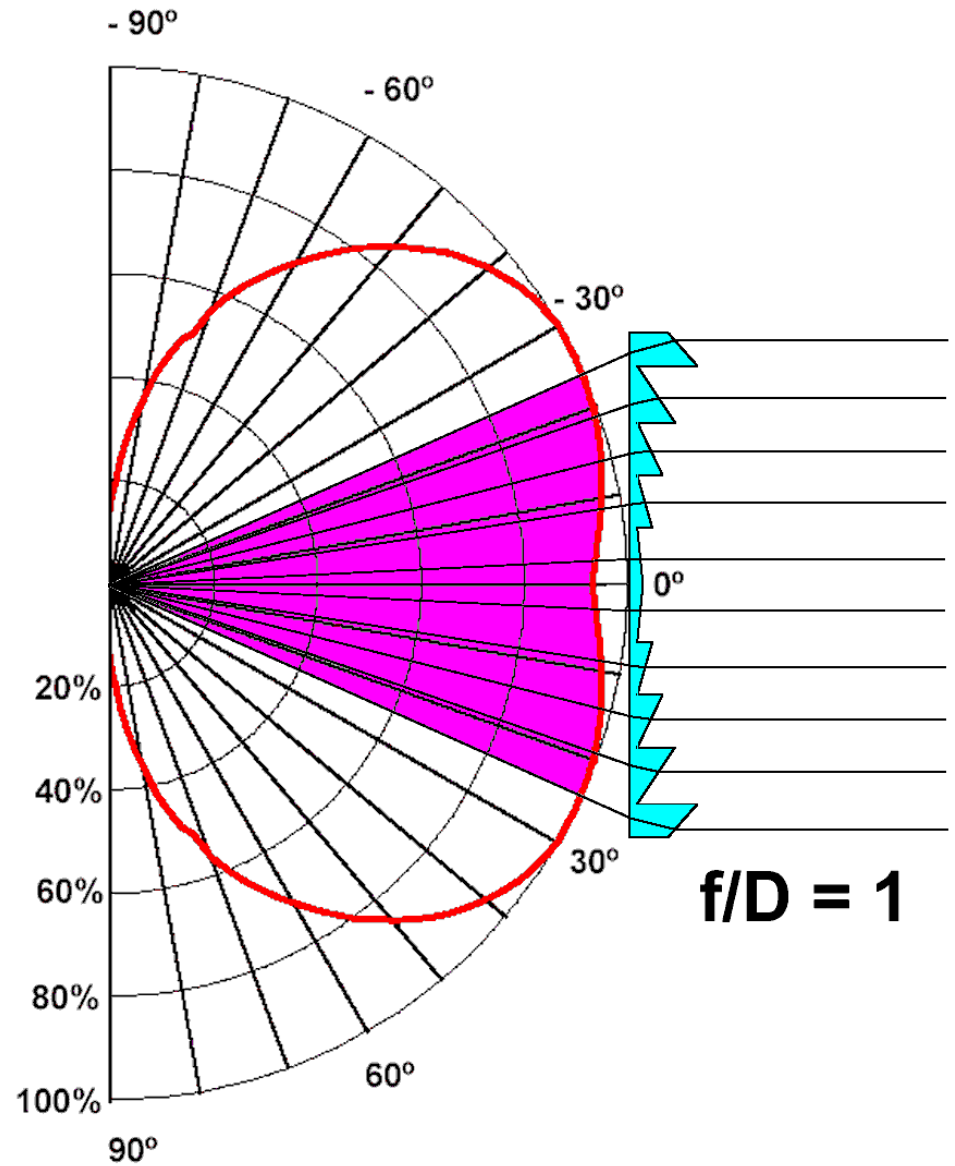
Receiver optics



Blur spot diameter $\sim D / 1000$ for Fresnel lenses

LED illuminating a Fresnel Lens

Much of the emitted light does not illuminate the Fresnel lens and is wasted.



G8AGN 3-D Vector Optical Ray Tracer

Available ray files are:

- 1... Luxeon K2 100k rays
- 2... Luxeon K2 500k rays
- 3... Golden Dragon 100k rays
- 4... Golden Dragon 500k rays
- 5... Golden Dragon 5M rays

Enter choice 5

Available ray traces are:

- 1..... Source alone seen at plane of Fresnel lens
- 2..... Source with secondary optics seen at plane of Fresnel lens
- 3..... Source with secondary and Fresnel at 100m
- 4..... Source with secondary and Fresnel at 1km
- 5..... Source with secondary and Fresnel at 10km

Enter choice 2

Choose secondary lens type:

- 1.... #8739 29mm diam, 28mm f.l.
- 2.... #10016 30mm diam, 26mm f.l.
- 3.... #4056 28.4mm diam, 28mm f.l.
- 4.... BC test

Choose : 3

Enter spacing between source and Fresnel lens 335.

Enter spacing between source and PMN lens 11.

Max hits in a pixel = 53

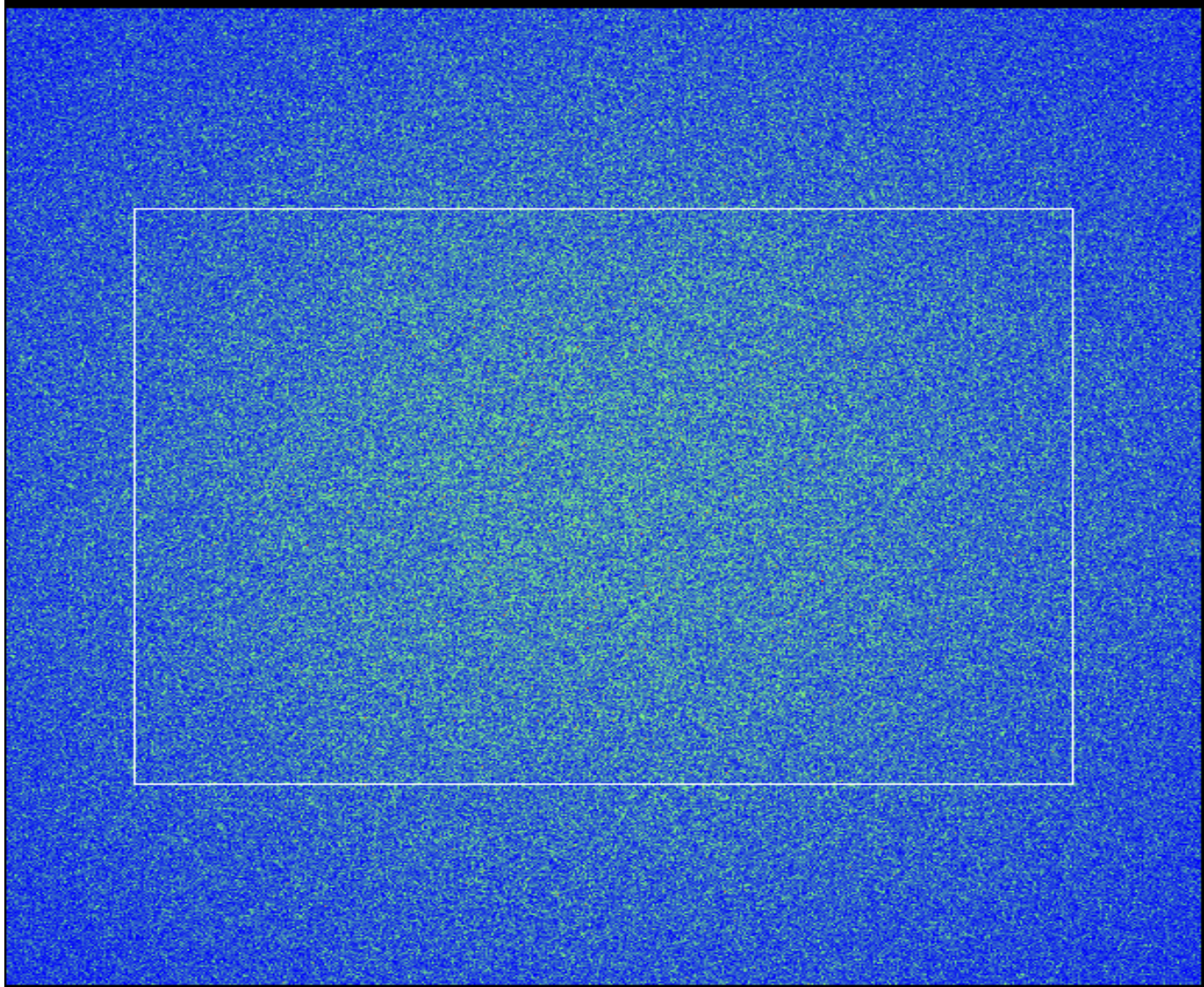
5000000 records read from ray file

Rays falling on Fresnel : 3276678

TRACE FINISHED

Press RETURN to close window . . .

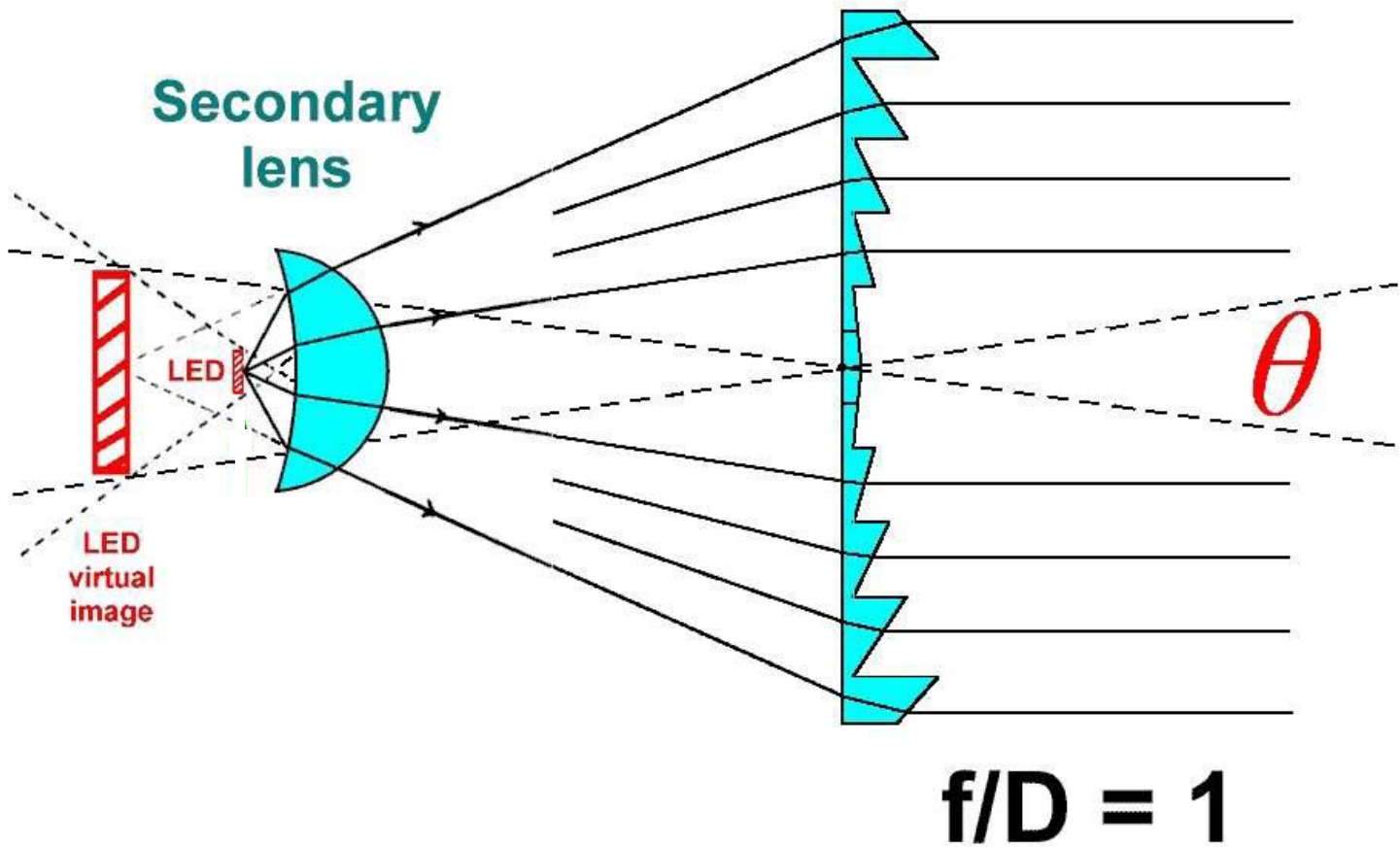
Golden Dragon 5M rays, Fresnel at 335mm, No PMN,
686618 rays hit Fresnel (13.7%)



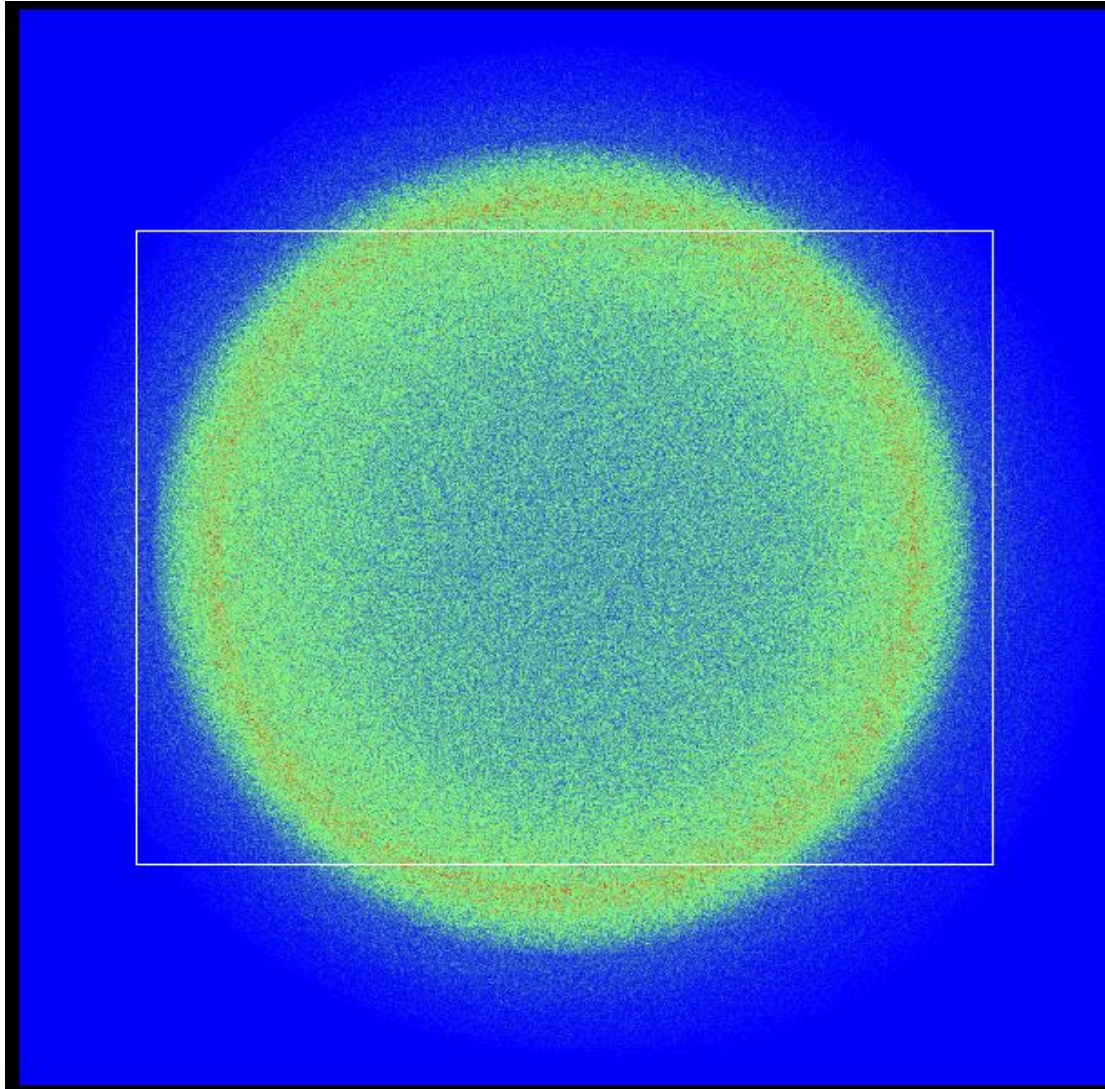
0.4 m

Using a secondary lens

Better lens illumination but greater beam divergence

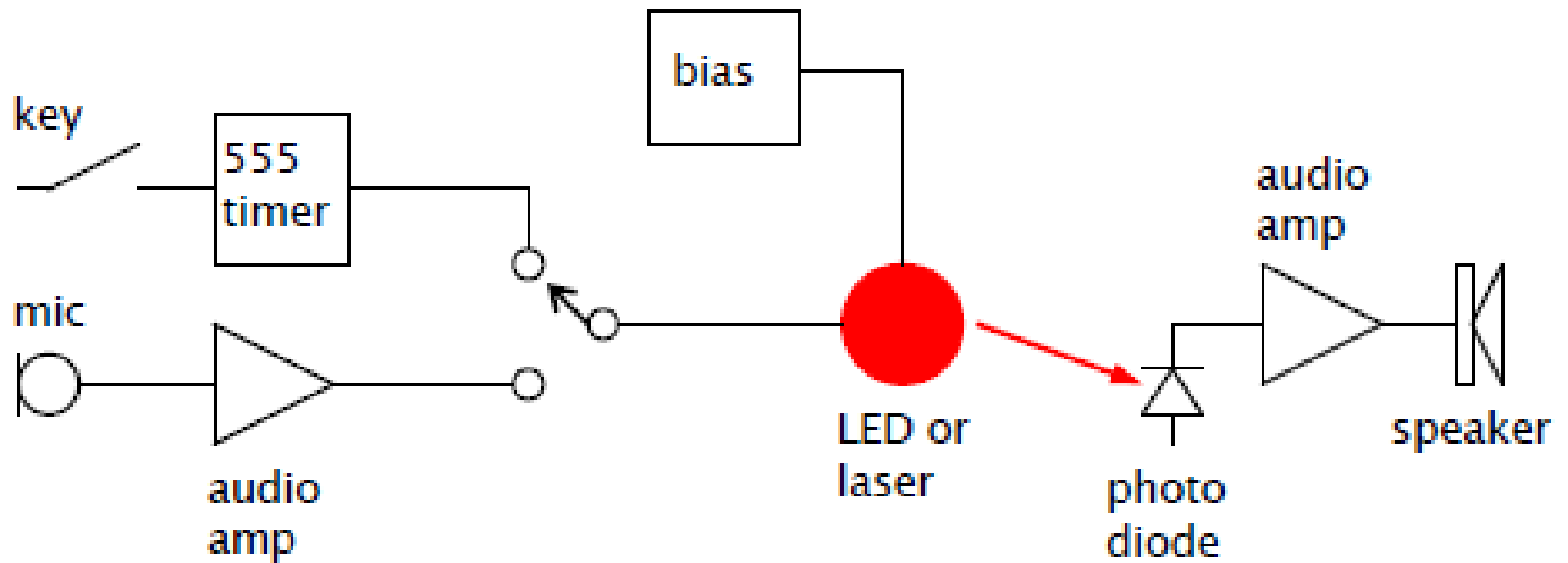


Golden Dragon 5M rays. #4056 lens at 11 mm, Fresnel at 335 mm
3108555 rays hit Fresnel (65.5%)



0.4m

Basic optical Tx and Rx for MCW or AM



G8AGN's optical transmitter and receiver

- Based on circuits developed by Clint, KA7OEI
- Tx uses Luminus *PhlatLight* LED 5 watt optical output, AM, MCW or data modulation
- Rx uses BPW34 photodiode and AF amplifier
- Separate Tx and Rx lenses (A4 page magnifiers)
- Solid construction and heavy duty tripod
- Telescopic sight for alignment

General view of G8AGN's optical transceiver

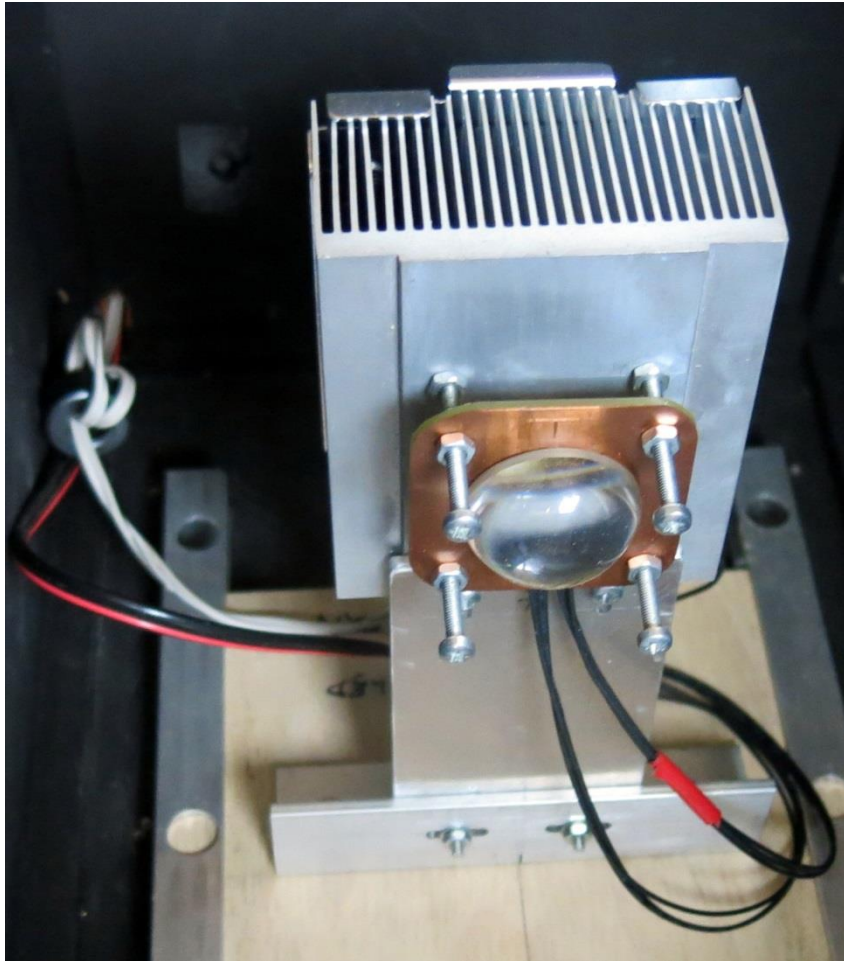
This photo was taken near Harpswell, Lincs
24 Nov 2010



G8AGN's optical transceiver – front view showing Fresnel lenses

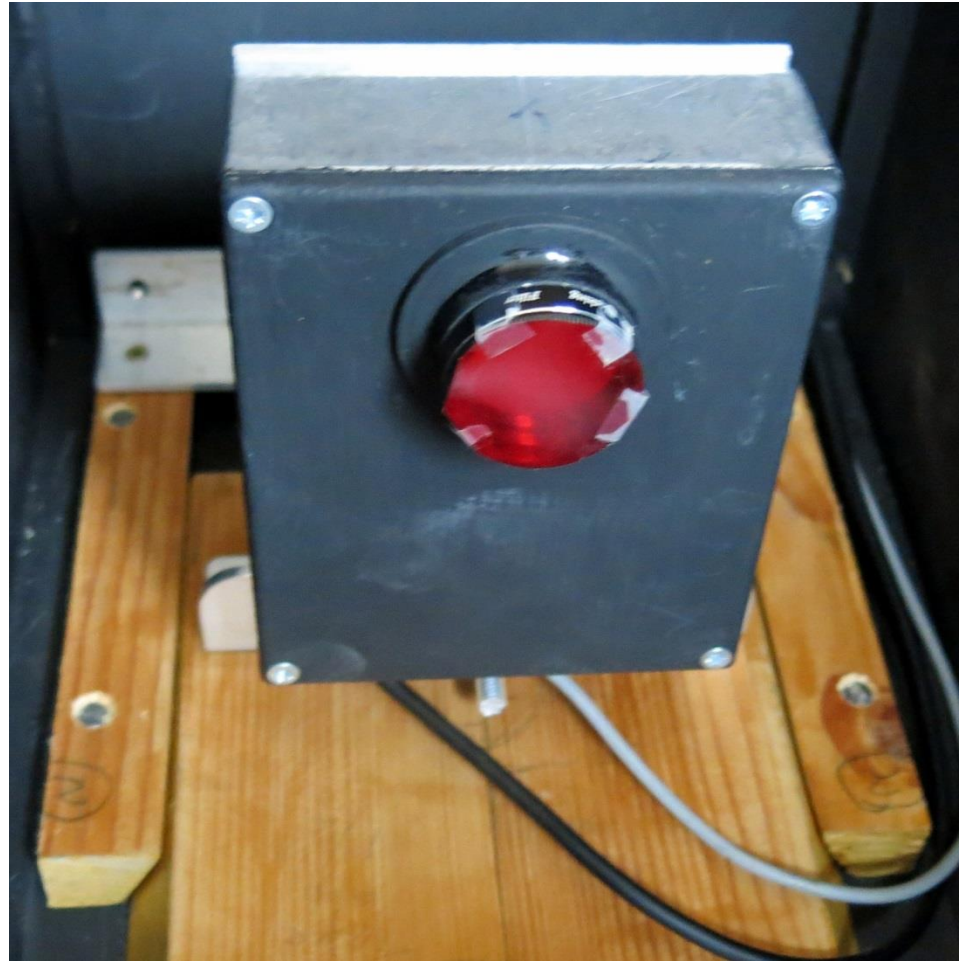


G8AGN optical transceiver – mounting of LED Tx and Rx front-end



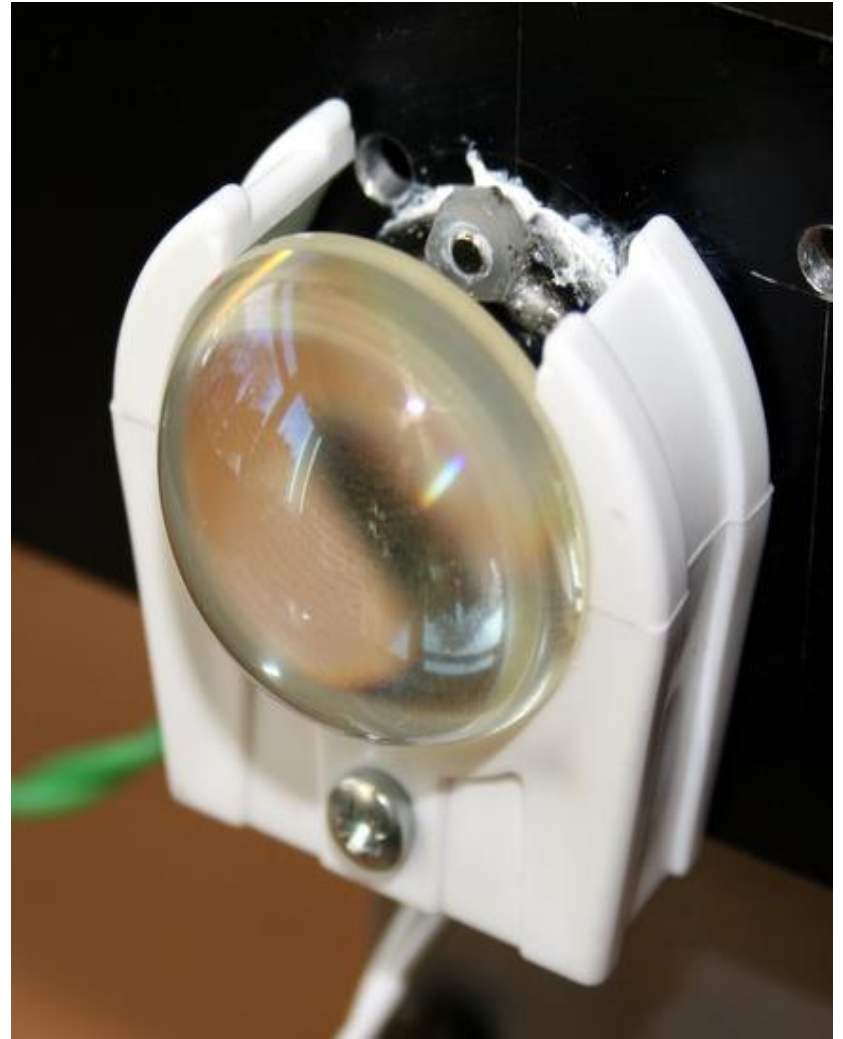
Tx

Fan-assisted cooling of LED



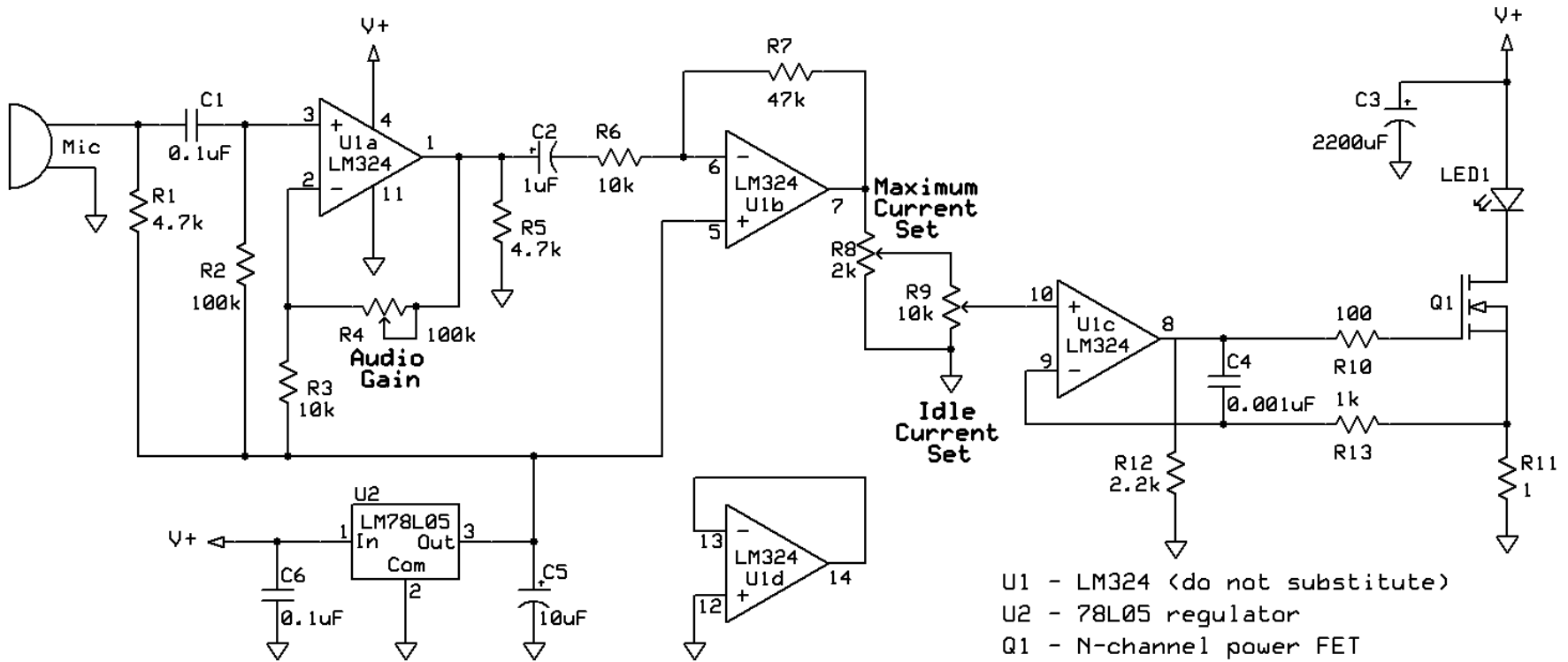
Rx

Red filter over photo-diode



Luxeon III mounted on heat-sink and fitted with PMN secondary lens

KA70EI current-sink modulator for high power LED



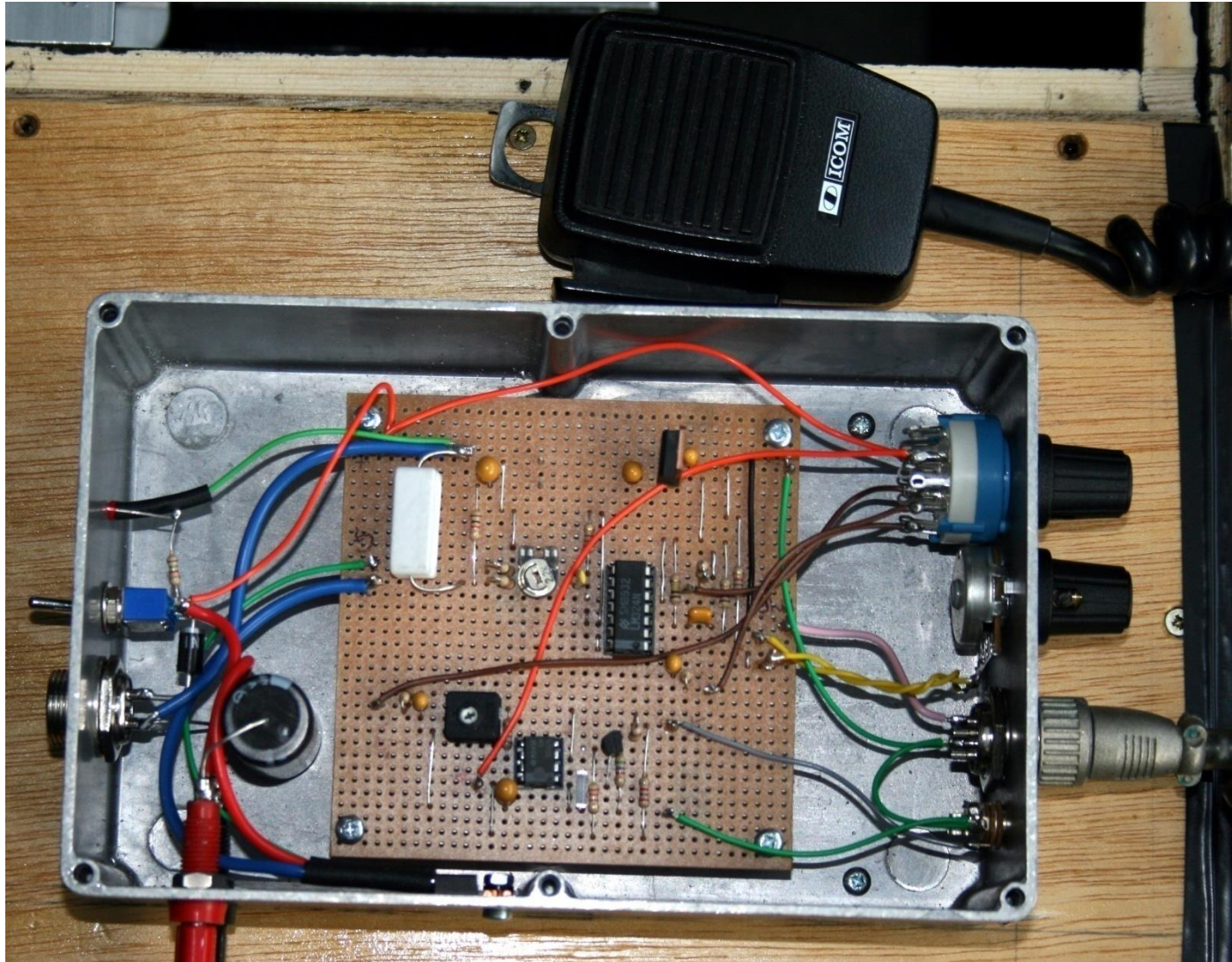
- U1 - LM324 (do not substitute)
- U2 - 78L05 regulator
- Q1 - N-channel power FET
- LED1 - High-power LED
- Mic - Electret Microphone

Simple high-current, adjustable LED modulator

Ver. 1.03b KA70EI

Light output of a LED is proportional to current, not voltage.

KA70EI linear modulator for high power LED



KA70EI optical receiver front-end

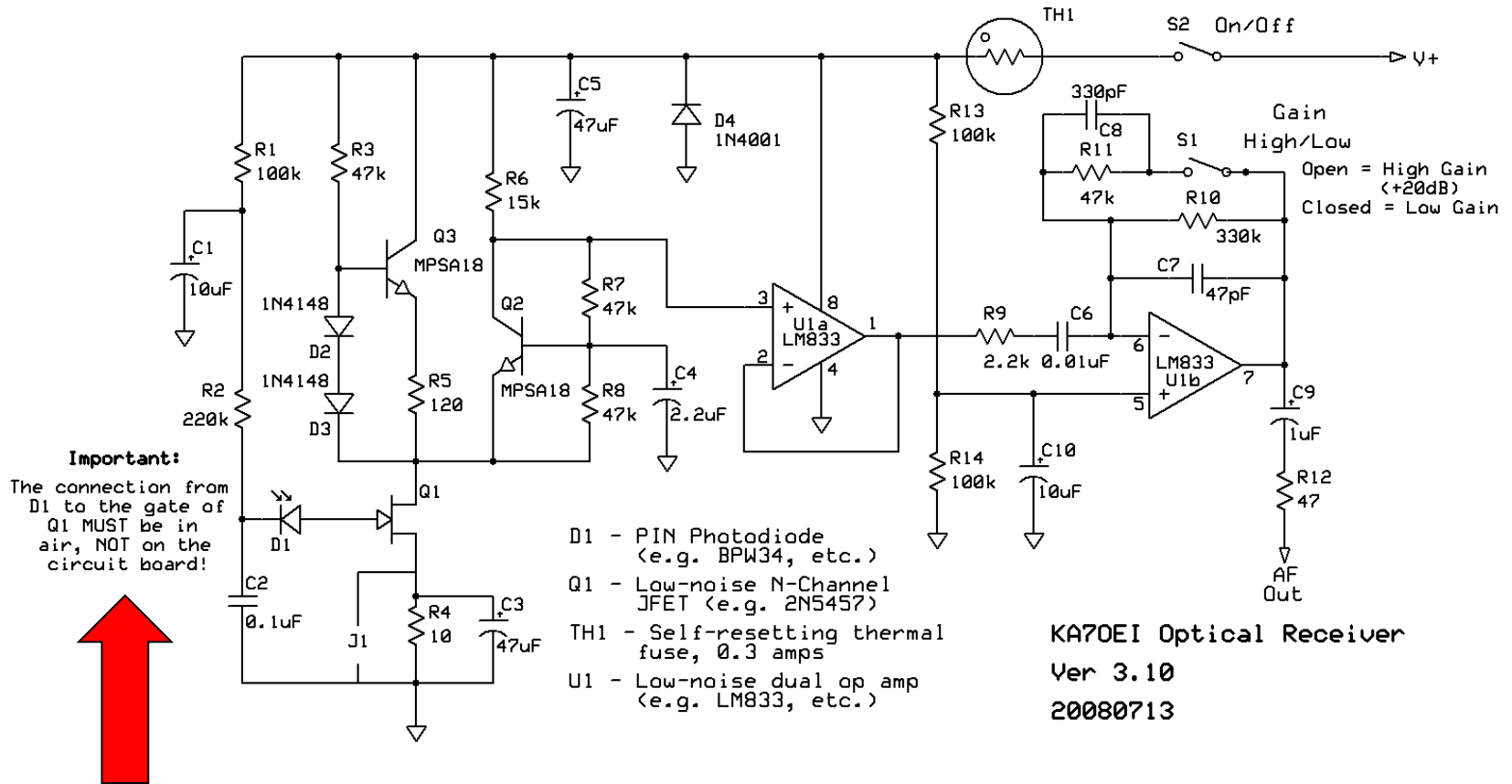


Photo-diode reverse-biased to reduce capacitance and improve bandwidth

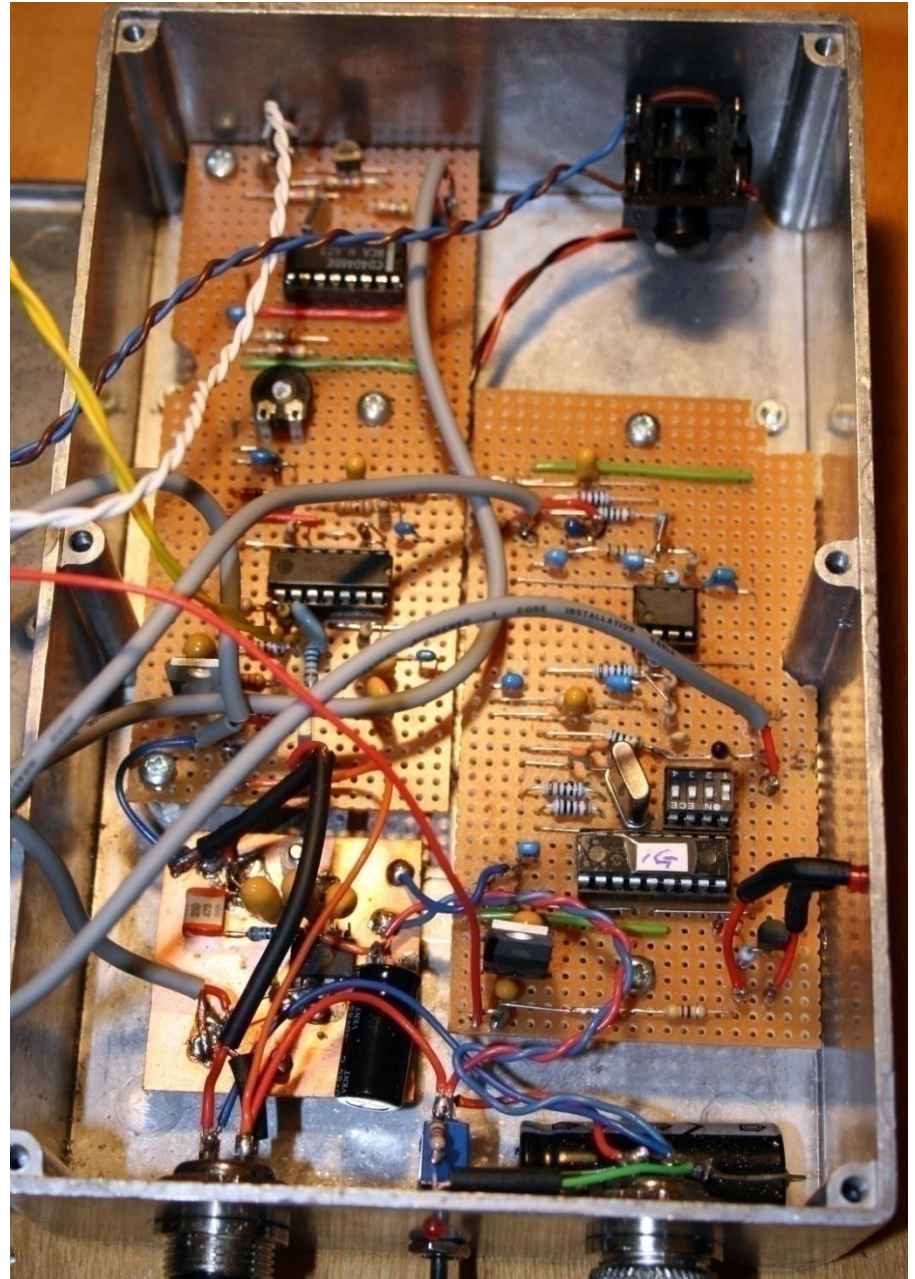
G8AGN optical Rx

This box is fed from a front-end comprising a photodiode and a low-noise pre-amplifier.

Shown are

- Audible S meter
- **PIC comb filter for removing 50 Hz and harmonics**
- Audio amplifier

All circuits due to KA7OEI



Optical receiver testing using a photon tube

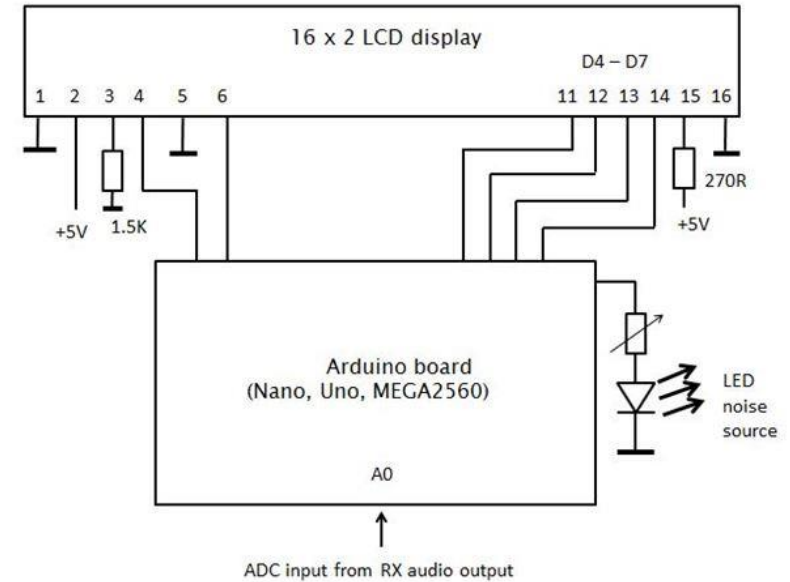
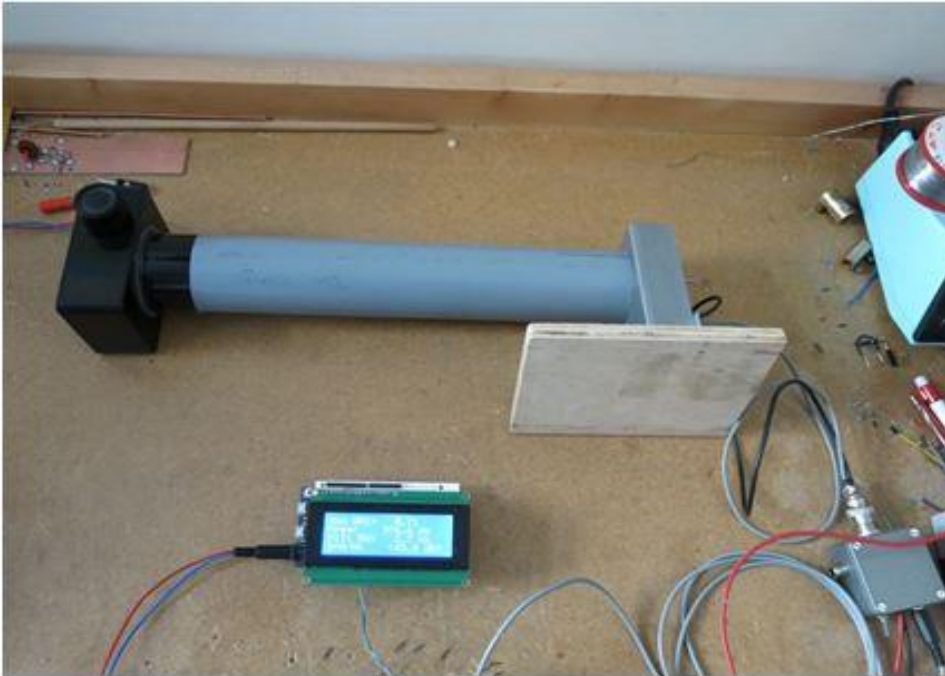


FIGURE 1: Hardware for Arduino S/N meter



In practice the photon tube is at least 2m long and lined with black flock paper to eliminate wall reflections

Optical target with light sensor



Optical target when illuminated
by G8AGN's Tx situated 100m
away

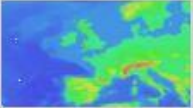


Path planning

Pocklington to Roper Hill – 87km

Radio path estimation (c) M.J. Willis
Help About

Input Data | Terrain Map | Site Database | Radio Refractivity Data | Tabulated Data | Zone Map | Mechanisms | Area Coverage



Calculate

Input Format

- Degrees Minutes Seconds
- Decimal Degrees
- OSGB Grid Reference
- Maidenhead Locator

Locations

53 56 47.64 N 0 43 6.61 W

53 21 25.83 N 1 35 0.45 W

Range km

Bearing deg

Radio Parameters

Frequency (GHz)

TX Mast Height (m)

RX Mast Height (m)

Humidity (%)

Temperature (C)

Pressure (mB)

Percentage

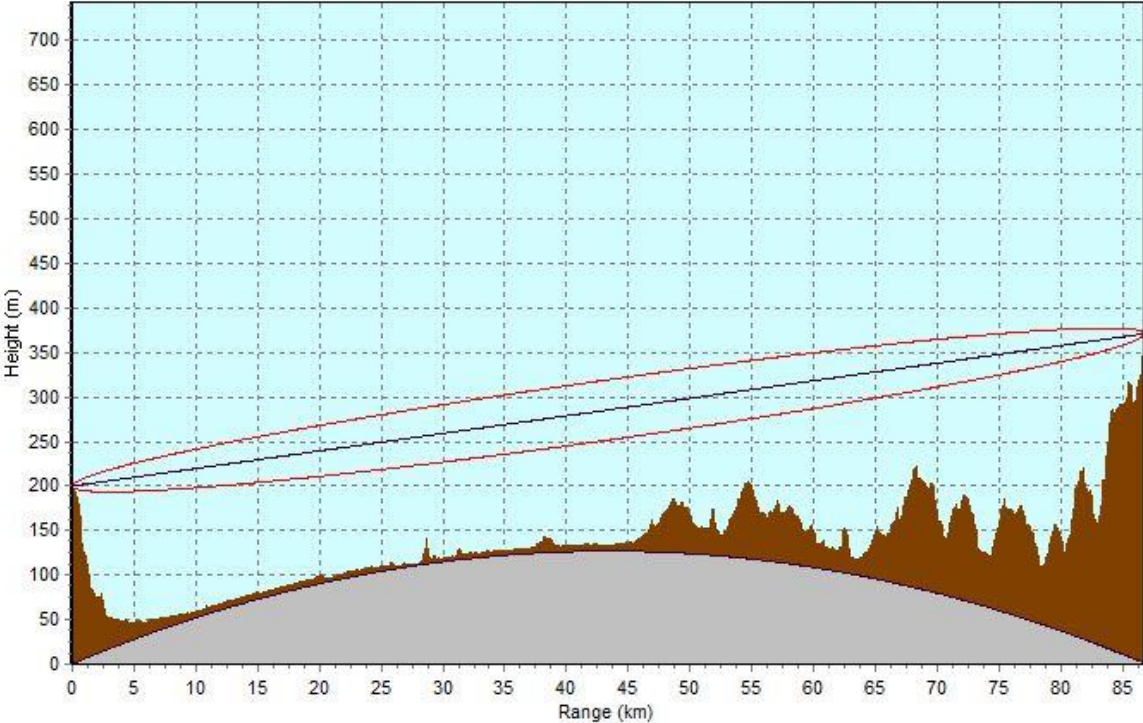
K factor Fixed

Results

RX ASL 378 m
TX ASL 198 m
Line of sight loss 147 dB

Gaseous Loss 0.7 dB
Diffraction path loss 155 dB
Troposcatter path loss 200 dB
Ducting path loss 244 dB

Overall path loss 155 dB



Height (m)

Range (km)

Profile Control

- Flat earth plot
- metric
- imperial
- Interpolate

Making an optical contact

- Pointing requires good Az/EI control
LEDs much easier due to wide beam
- Aligned rifle scope or viewing through receiver optics
- Target must be visible unless precise Az/EI calibration and data is available
- Liaison radio link for coordination and audio pointing feedback
- Atmospheric scintillation can be a problem

G8AGN/P's 0.5W Tx as seen by G0EWN/P over an 87km path



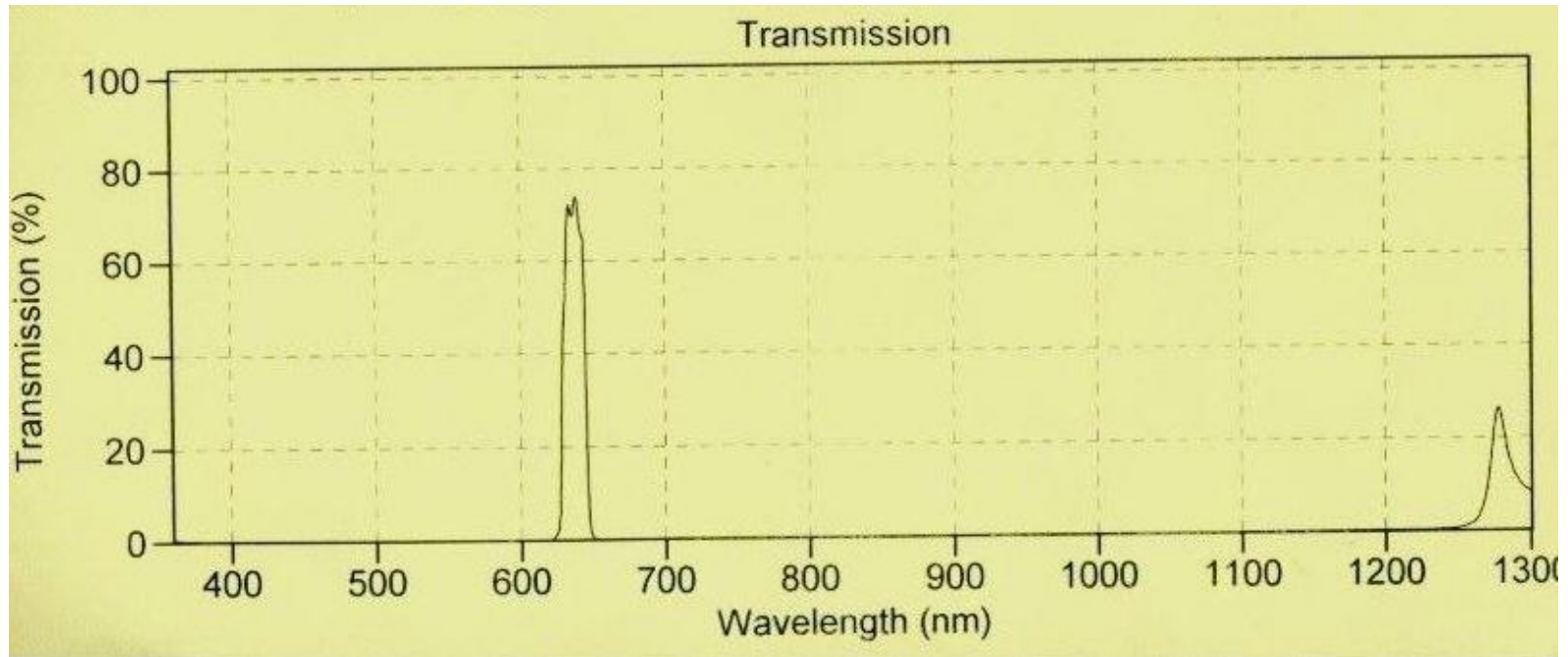
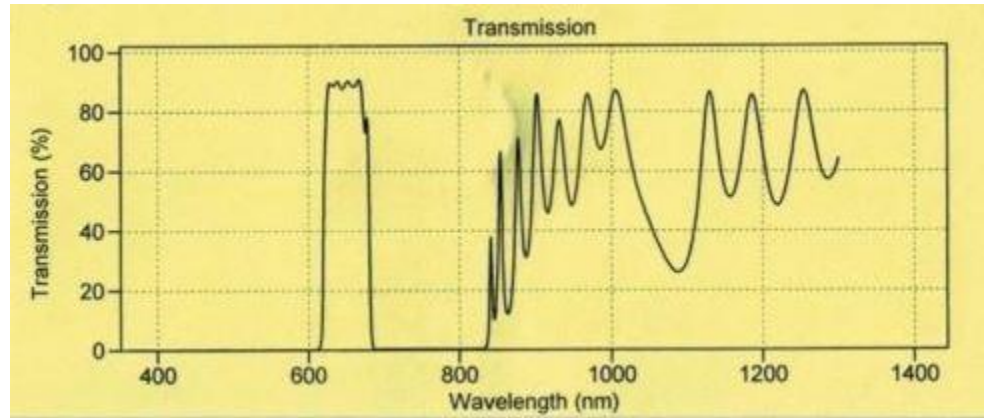
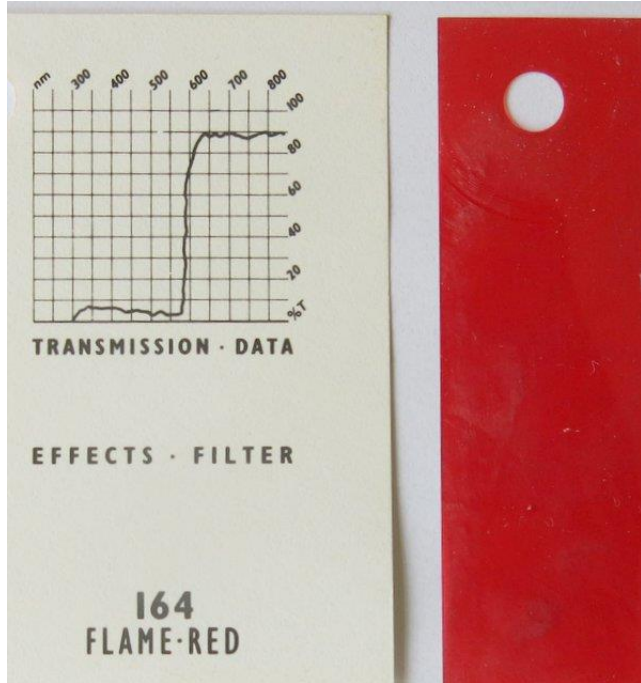
Note the QRM

Contact made on 8 Jan 2011

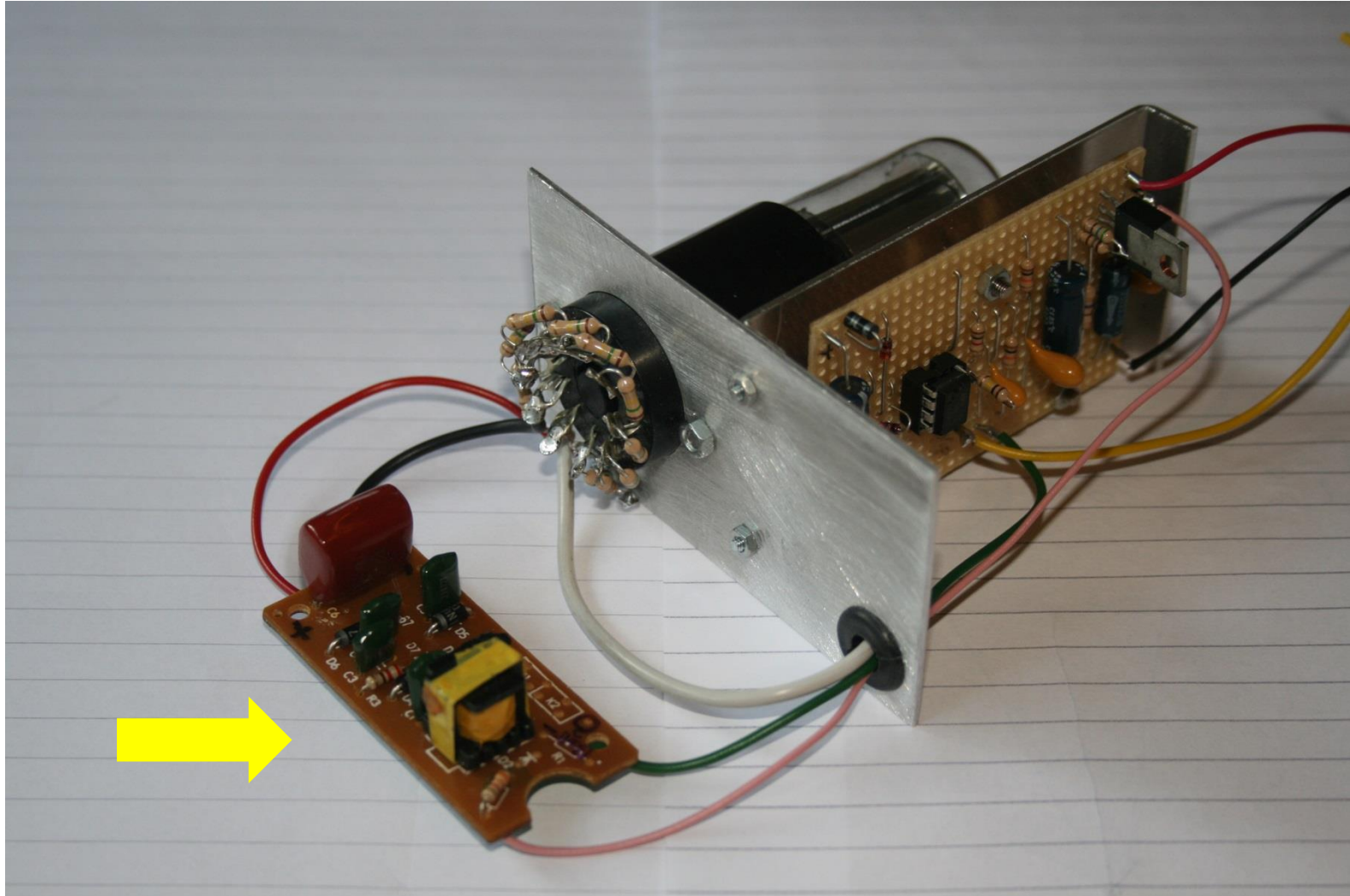
Optical DX records

Heliograph	293km	US Army	1894
Line of sight (laser)	189km	W6QYY, W6OP	1963
Line of sight (LED)	278km	KA7OEl group	2007
Cloud Bounce (LED)	288km (1 way)	VK7MO group	2009
UK (LED)	129km	G8AGN, G0EWN	2012
Daylight (LED)	83km	G8AGN, G0RPH	2013

Optical Filters

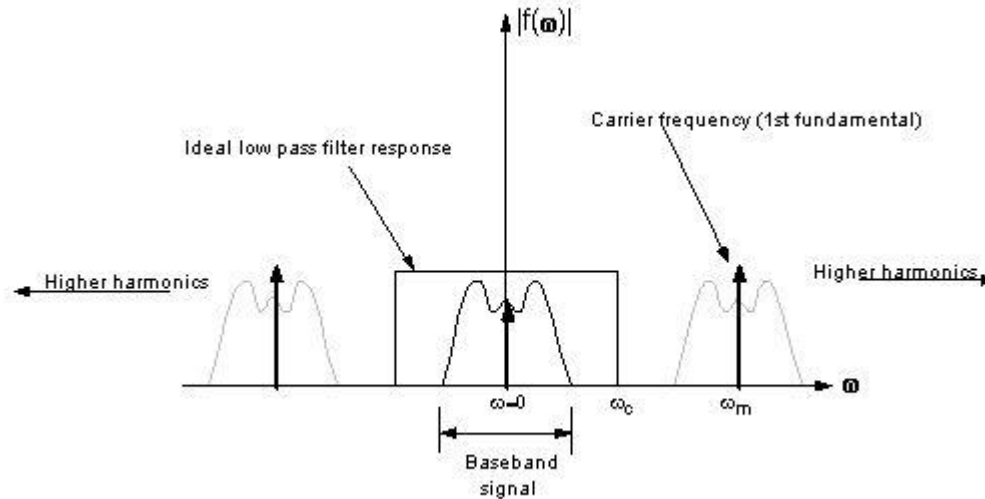


UV receiver using 931A photo-multiplier tube

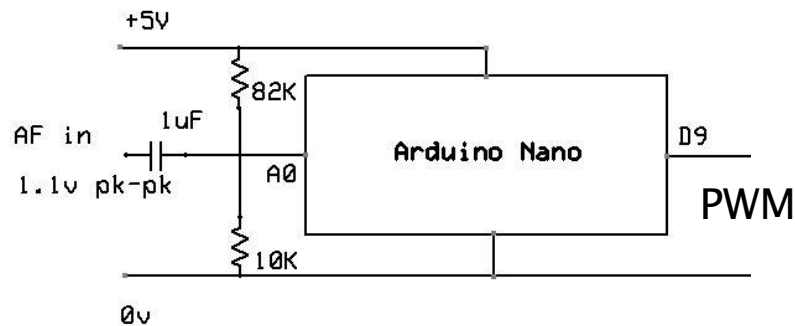


High voltage PSU for 931A taken from low-cost “bug zapper”

LED pulse width modulation



Default PWM frequency = 490 Hz
Reconfigure Timer 1 to increase this

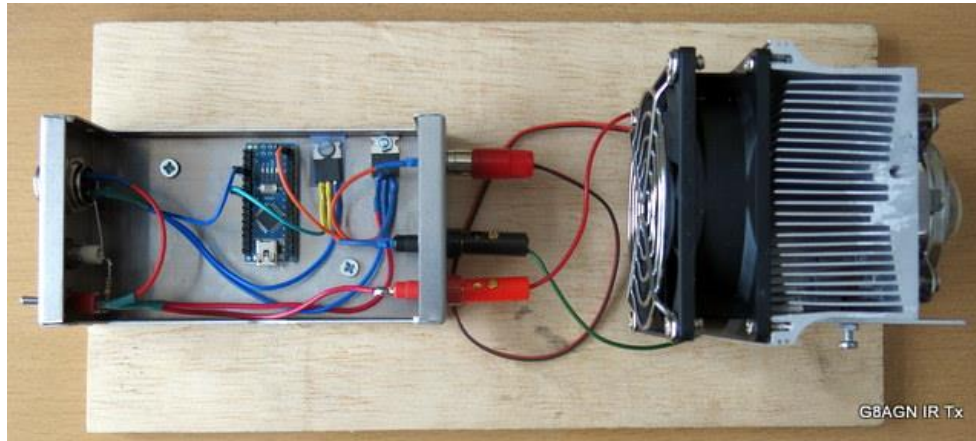


Clock frequency = 16 MHz
Number of bits = 10 = 1024 levels
Sampling frequency = $16000/1.024$
= 15.625 kHz

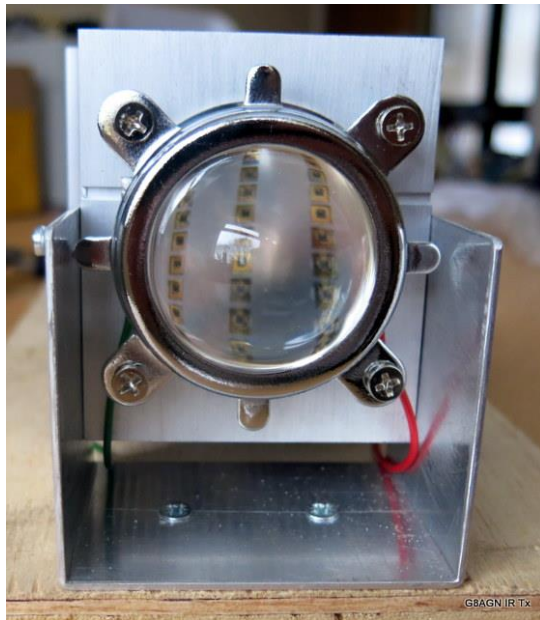
wiki.openmusiclabs.com/wiki/PWMDAC

Use Arduino internal 1.1 V voltage reference

SMT-Hell modulator for LED



Arduino Nano version (G8AGN)



A more versatile approach is the PIC based

SCRIBER module developed by G8HAJ

Details in Scatterpoint Feb 2019

Where to Find More Information

www.ka7oei.com

www.k3pgp.org/

www.barry-chambers.staff.shef.ac.uk/LED_files/led.html

www.pageperso.aol.fr/YvesF1AVY/UKINDEX.html

www.ham-radio/sbms/sd

<http://g8haj.uk/>

Yahoo Group on UKNanowaves

Yahoo Group on Optical DX

www.surplussed.com (for lenses)

www.ebaystores.co.uk/BJOMEJAG-EBUYER-STORE (for optical filters)

“An Arduino-based receiver noise figure alignment aid”, Barry Chambers G8AGN, RadCom, 2014

“Exploits in lightwave communications”, Stuart Wisher G8CYW, Practical Wireless, 2013



Thank you for listening.

Any questions ?